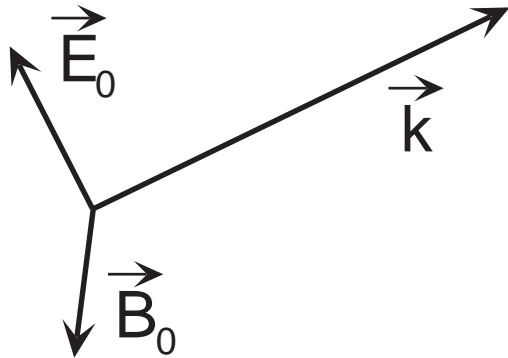


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8.044 Statistical Physics I  
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Thermal Radiation Radiation in thermal equilibrium  
with its surroundings



$$\vec{E} = \vec{E}_0 e^{i(\vec{k} \cdot \vec{r} - \omega t)}$$

$$\vec{B} = \vec{B}_0 e^{i(\vec{k} \cdot \vec{r} - \omega t)}$$

$$\omega = c|\vec{k}|$$

$$\vec{B}_0 = \hat{\mathbf{1}}_k \times \vec{E}_0 / c$$

Time average energy density

$$\bar{u} = \frac{1}{2}\epsilon_0|\vec{E}_0|^2$$

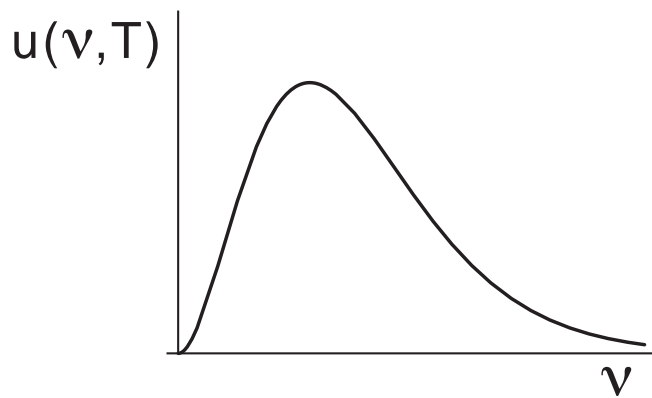
Time average energy flux

$$\vec{j}_E = (c\bar{u})\vec{1}_k$$

Time average pressure ( $\perp$  to  $\vec{k}$ )

$$P = \bar{u}$$

Thermal radiation has a continuous distribution of frequencies.

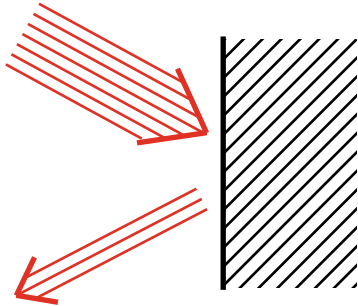


Peaks near  $h\nu = 3k_B T$

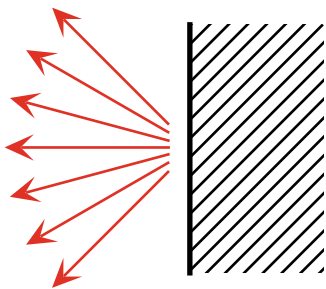
$(h/k_B \sim 5 \times 10^{-11} \text{ K-sec})$

Spectral Region	$\nu$ (Hz)	$T$ (K)	Thermal Rad.
Radio	$10^6$	$5 \times 10^{-5}$	
Microwave	$10^{10}$	0.5	cosmic background
Infrared	$10^{13}$	$5 \times 10^2$	room temp.
Visible	$\frac{1}{2} \times 10^{15}$	$2 \times 10^4$	sun's surface
Ultraviolet	$10^{16}$	$5 \times 10^5$	
X ray	$10^{18}$	$5 \times 10^7$	black holes
$\gamma$ ray	$10^{21}$	$5 \times 10^{10}$	

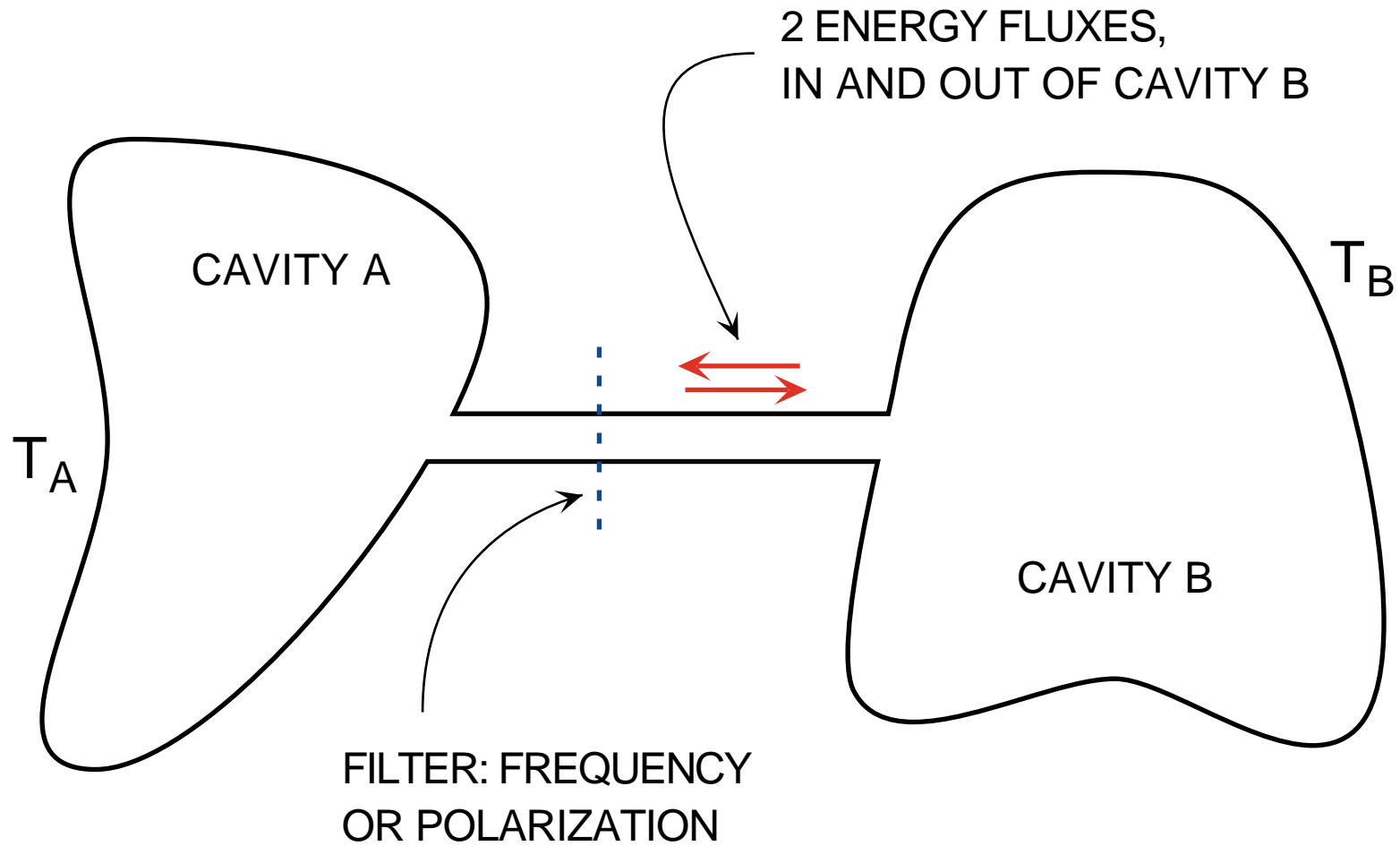
ABSORPTIVITY  $\alpha(\nu, T) \equiv \left\langle \frac{\text{ENERGY ABSORBED}}{\text{ENERGY INCIDENT}} \right\rangle$   
ISOTROPIC



EMISSIVE POWER  $e(\nu, T) \equiv \left\langle \frac{\text{ENERGY EMITTED}}{\text{AREA}} \right\rangle$   
ISOTROPIC



# THERMAL RADIATION: PROPERTIES



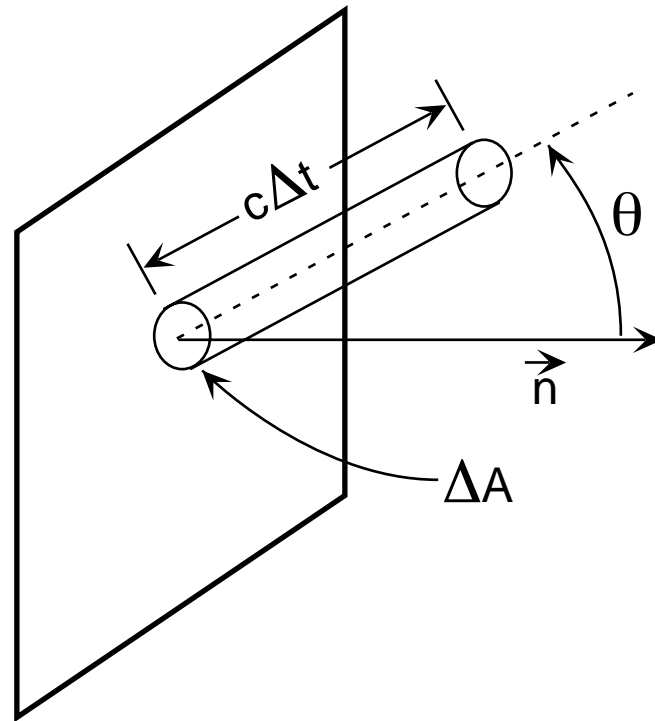
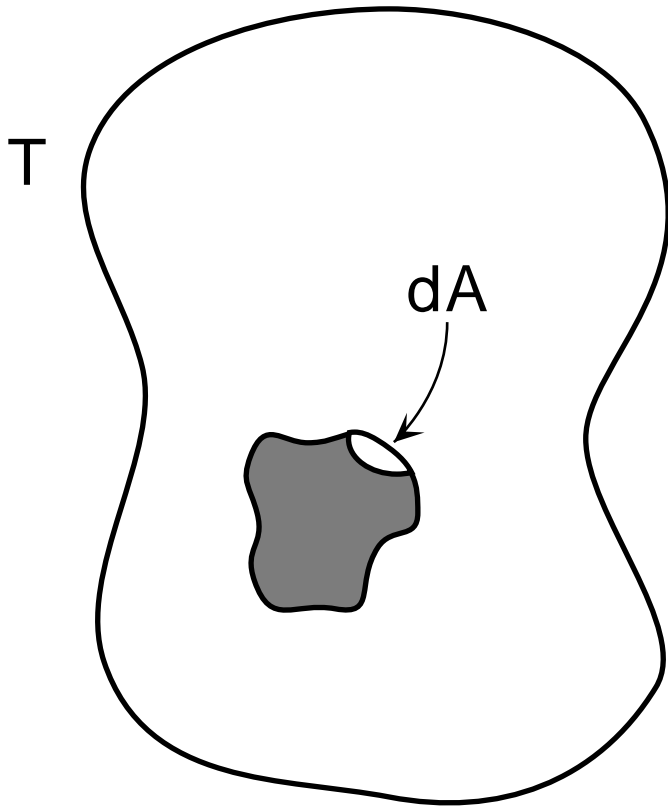
ASSUME  $T_A = T_B$  AND THERMAL EQUILIBRIUM

## CONCLUSIONS:

- $u(\nu, T)$  is independent of shape and wall material
- $u(\nu, T)$  is isotropic
- $u(\nu, T)$  is unpolarized

CONSIDER AN OBJECT IN THE CAVITY,  
IN THERMAL EQUILIBRIUM

COMPUTE THE ENERGY FLUX





$$\begin{aligned}
\Delta E &= \int (E \text{ in cylinder}) p(\theta, \phi) d\theta d\phi \\
&= \int (u \Delta A \cos \theta c \Delta t) \left( \frac{\sin \theta}{2} \frac{1}{2\pi} \right) d\theta d\phi \\
&= c u \Delta A \Delta t \underbrace{\int_0^{\pi/2} \frac{\cos \theta \sin \theta}{2} d\theta}_{1/4} \underbrace{\int_0^{2\pi} \frac{1}{2\pi} d\phi}_1
\end{aligned}$$

$$\Rightarrow \text{energy flux onto } dA = \underline{\underline{\frac{1}{4} c u(\nu, T)}}$$

## Momentum Flux

Plane wave momentum density  $\vec{p} = \frac{u}{c} \vec{1}_k$

$|\Delta p| = 2|p_{\perp}|$  since  $\vec{p}_{\perp \text{ in}} = -\vec{p}_{\perp \text{ out}}$

$$\begin{aligned}
|\Delta p|_\nu &= \int \left( \frac{2 \cos \theta}{c} \right) (E \text{ in cylinder}) p(\theta, \phi) d\theta d\phi \\
&= u(\nu, T) \Delta A \Delta t \underbrace{\int_0^{\pi/2} \cos^2 \theta \sin \theta d\theta}_{1/3} \underbrace{\int_0^{2\pi} \frac{1}{2\pi} d\phi}_1 \\
&= \frac{1}{3} u(\nu, T) \Delta A \Delta t
\end{aligned}$$

$$\Rightarrow P(T) = \underline{\frac{1}{3} \int_0^\infty u(\nu, T) d\nu}$$

Apply detailed balance to the object in the cavity.

$$E_{\text{out}} = E_{\text{in}}$$

$$e dA = \alpha \left( \frac{1}{4} c u(\nu, T) \right) dA$$

$$\Rightarrow \frac{e(\nu, T)}{\alpha(\nu, T)} = \frac{1}{4} c u(\nu, T)$$

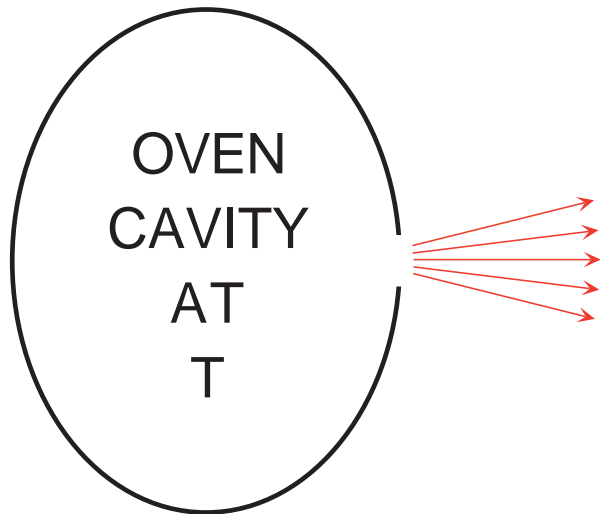
This ratio has a universal form for all materials.

The result is known as KIRCHOFF'S LAW.

# Black Body Radiation

If  $\alpha \equiv 1 \equiv$  “Black”

$$\text{Then } e(\nu, T) = \frac{1}{4}c u(\nu, T)$$



Measure  $e(\nu, T)$   
and obtain  $u(\nu, T)$